DEEPWAVE: Meteoric Plasma Dynamics Radar System

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- 1. Radar facility for Meteor plasma probing
- Measuring wind speeds and turbulent structure 80 – 110 km
- 4. Optical probing from Mount John.







Kaikorete spit – looking west



Meteor Orbit Radar transmitter 26 MHz: Looking North-East



Meteor Orbit Radar transmitter 26 MHz: Looking North



PROGRAMME:

- 1. Atmospheric wind patterns 80 110 km
- 2. Small scale effects: wind-shear and turbulence

Phase behaviour of radar scattering from meteoric plasma: Sampling rate of echoes ~ 1 per second

Plasma location resolution



Radial phase signature & Wind components





Phase records to provide radial wind speed

Echo amplitude

Echo Phases



Echo amplitude

Echo Phases



Radar pulses



Target receding

Radar pulses

Echo Phases Echo amplitude 50 F 250 F Amplitude (HOME) mhanahanahan 200 Mark Marken M H) 30 20 10 0 Mm/MW/W how when the PHASE 1 150 E 100 50 01 250 50 100 150 200 0 50 100 150 200 250 0 300 E 250 Radar pulses ∼ 200¥ PHASE 150 Ĕ 100 E 50 E 0 250 0 50 100 150 200 300 E 250 È m 200 PHASE 150 100 50 ~ ~ ~ 0 - 0 Phase increase with time-Radar pulses Target approaching

Echo amplitude

Echo Phases





MEASURING WIND SHEAR

2. Change of phase at one antenna

 $\Delta R = 0.5 R \theta^{2}$ $\theta \sim S \sin Z$ $\lambda = 7m, R \sim 200 \text{ km}, Z \sim 60^{\circ}$ $S \sim 1 \text{ ms}^{-1} \text{ km}^{-1}$

 $\Delta R = 0.5 \ 200 \ 10^3 \ (10^{-3})^2 \ . \ 0.8$

 $\Delta R \sim 0.1 \text{ m s}^{-1}$ Phase change twice this ~ 7 deg per second ~ 0.018 degrees per radar pulse

Sign of shear from sign of elevation change

TX RX

R

 $R - \Delta R$

System **uses Circular Polarisation** – reduces any effects of plasma resonance distorting the phase record

Each element: crossed dipole with reflector and fed via $\lambda/4$ phase section



ground









View looking east from height ~300m



View looking south from height~ 100m

A - I - I - I - I - I - I - I - I - A

View looking south from ~ 200m







Multi-wavelength All-Sky Imager Steve Smith Bo

Fabry-Perot Interferometer



All-Sky Imager fish-eye lens and dome.



All-Sky Imager fish-eye lens and dome.

Two periscopes of FPI background



Multi-wavelength All-Sky Imager:

- OH 800 nm 87 km
- O (¹S) 577 nm 96
- Na (²S) 589 nm 90
- O (1D) 630 nm 250

All-Sky sampling Provides: patterns of emission ~ 0.1 deg' Horizontal speeds ; accurate positions

Mount John four band spectral Imager



OH

Na

O

Steve Smith Boston Univ.

Mount John All-Sky 557.7 nm O¹S



Maps of wave structure at heights ~ 95 km Steve Smith Boston Univ.





Steve Smith Boston Univ.

Fabry-Perot Interferometer:

OH 800 nm 87 km O2 860 nm 94 O (¹S) 577 nm O (1D) 630 nm

Samples N, E, S, W at ~ 20 degrees and zenith ; approx positions Provides - Horizontal speeds and temperatures;







FPI



Support for DeepWave



Adrian McDonald, University of Canterbury

- SNOW WEB deployment in Southern Alps.
- CL51 Ceilometer deployment in Southern Alps or Canterbury Plains.
- Hosting radiosonde receiver station on Rutherford building.
- Support for balloon launching (some financial commitment will be required if this is needed significantly).
- Potential office space if required.

SNOW WEB: Hardware development





Our capability in hardware development has allowed us to develop an expandable Zigbee-enabled environmental monitoring system.

SNOW WEB allows the possibility of rapidly expanding the density of observations of atmospheric (wind velocity, temperature, pressure and humidity) and cryospheric (ice velocity and snow depth) properties over broad regions.

SNOW WEB: Model Validation



model

observations

Mesoscale model output (left) and SNOW WEB observations in Antarctica from the 2011/2012 deployment show the model misrepresents the impact of topography on the flow.

SNOW WEB: A Movie





CL51 Vaisala Ceilometer





CL51 Vaisla Ceilomter can be used to examine cloud and boundary layer heights by examining backscatter.



